



**DAOISM**

## Neural network-based emulation of ISM models

P. Palud, L. Einig, F. Le Petit, E. Bron, P. Chainais, J. Chanussot  
and the ORION-B consortium



# Interstellar medium and numerical simulations

**Numerical simulations** are widely used in order to model the ISM and compare theory with observations.

# Interstellar medium and numerical simulations

**Numerical simulations** are widely used in order to model the ISM and compare theory with observations.

Computation time is often prohibitive for inference procedures.

# Interstellar medium and numerical simulations

**Numerical simulations** are widely used in order to model the ISM and compare theory with observations.

Computation time is often prohibitive for inference procedures.

## Usual solutions



# Interstellar medium and numerical simulations

**Numerical simulations** are widely used in order to model the ISM and compare theory with observations.

Computation time is often prohibitive for inference procedures.

## Usual solutions

- Interpolation methods:

- 1 Nearest point in grid (Sheffer+2011; Sheffer+2013; Joblin+2018)
- 2 SciPy interpolation (Wu+2018; Ramambason+2022)

# Interstellar medium and numerical simulations

**Numerical simulations** are widely used in order to model the ISM and compare theory with observations.

Computation time is often prohibitive for inference procedures.

## Usual solutions

### ■ Interpolation methods:

- 1 Nearest point in grid (Sheffer+2011; Sheffer+2013; Joblin+2018)
- 2 SciPy interpolation (Wu+2018; Ramambason+2022)

### ■ Regression-based approximations:

- 1 *k*-nearest neighbors (Smirnov-Pinchukov+2022)
- 2 Random forests (Bron+2021)
- 3 Neural networks (de Mijolla+2019; Holdship+2021; Grassi+2022)

# Interstellar medium and numerical simulations

**Numerical simulations** are widely used in order to model the ISM and compare theory with observations.

Computation time is often prohibitive for inference procedures.

## Usual solutions

### ■ Interpolation methods:

- 1 Nearest point in grid (Sheffer+2011; Sheffer+2013; Joblin+2018)
- 2 SciPy interpolation (Wu+2018; Ramambason+2022)

### ■ Regression-based approximations:

- 1  $k$ -nearest neighbors (Smirnov-Pinchukov+2022)
- 2 Random forests (Bron+2021)
- 3 Neural networks (de Mijolla+2019; Holdship+2021; Grassi+2022)  
→ Less complex, so faster and allow more training data

# A Meudon PDR approximation as a template

## The Meudon PDR code (Le Petit+2006)

- Emulates **photo-dissociation regions** (PDRs) at equilibrium.
- This version: **4** inputs  $\mapsto \sim$  **5 000** spectral lines intensities
- Execution time  $\sim$  **6 hours** and may yield **anomalies**.
- Predictions **directly comparable with observations**.

# A Meudon PDR approximation as a template

## The Meudon PDR code (Le Petit+2006)

- Emulates **photo-dissociation regions** (PDRs) at equilibrium.
- This version: **4** inputs  $\mapsto \sim$  **5 000** spectral lines intensities
- Execution time  $\sim$  **6 hours** and may yield **anomalies**.
- Predictions **directly comparable with observations**.

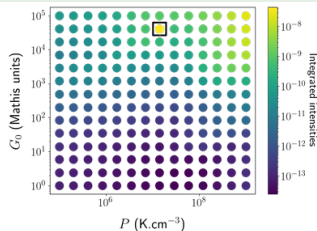
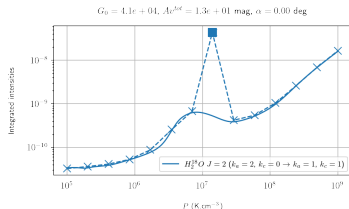
**Overall:** representative example of ISM numerical simulations.

# A Meudon PDR approximation as a template

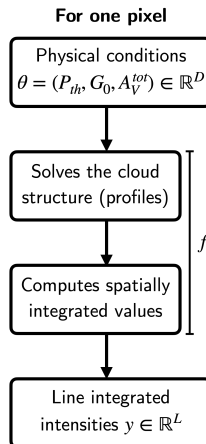
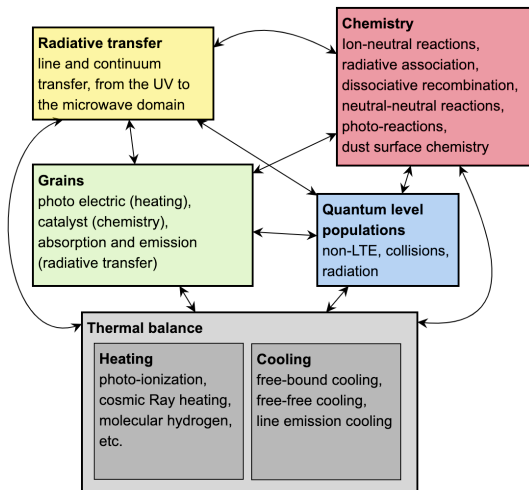
## The Meudon PDR code (Le Petit+2006)

- Emulates **photo-dissociation regions** (PDRs) at equilibrium.
- This version: **4** inputs  $\mapsto \sim$  **5 000** spectral lines intensities
- Execution time  $\sim$  **6 hours** and may yield **anomalies**.
- Predictions **directly comparable with observations**.

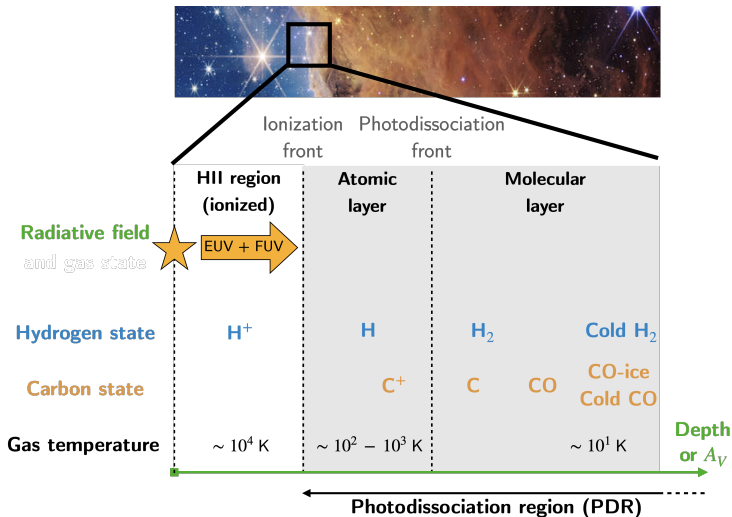
**Overall:** representative example of ISM numerical simulations.



# The Meudon PDR code

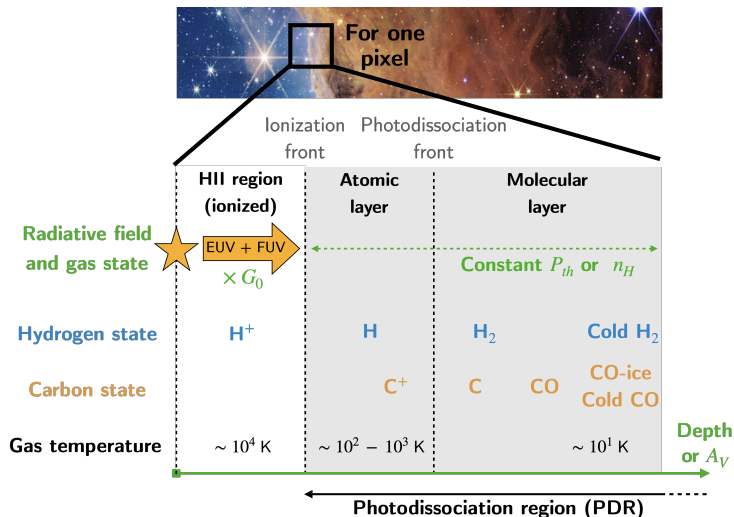


# The Meudon PDR code

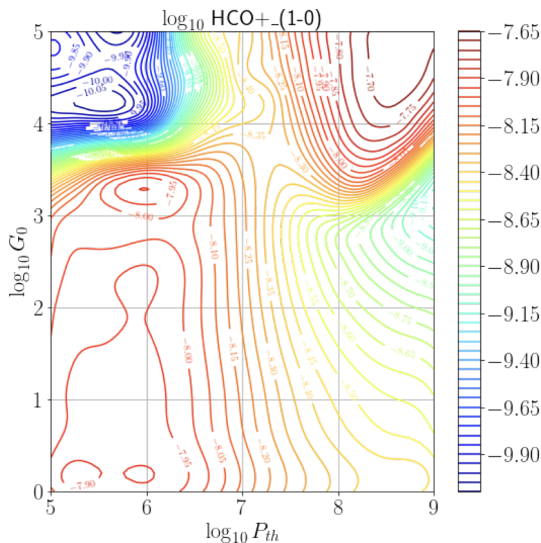




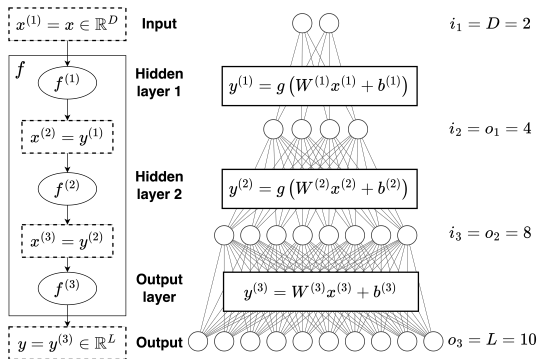
# The Meudon PDR code



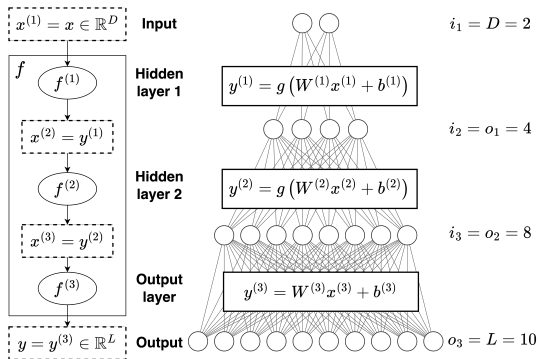
# The Meudon PDR code



# Evolutions of a standard multilayer perceptron



# Evolutions of a standard multilayer perceptron



## Proposition 1: polynomial expansion of inputs

Can help a network creating non-linearities. It has to be done **at execution** to ensure the network to be **fully differentiable**.

**Ex:**  $P_2(x_1, x_2, x_3) = (x_1, x_2, x_3, x_1^2, x_1x_2, x_1x_3, x_2^2, x_2x_3, x_3^2)$

# Evolutions of a standard multilayer perceptron

## Proposition 2: ignoring anomalies

**Anomalies**  $\neq$  well-modeled points with sensitive behavior!

- 1 Training with a **robust loss** (e.g., Cauchy) to detect badly reconstructed points.
- 2 Use physics knowledge to determine anomalies among them.
- 3 New training from scratch with a **masked non-robust loss function** (e.g., MSE), **ignoring the abnormal outputs**.

# Evolutions of a standard multilayer perceptron

## Proposition 2: ignoring anomalies

**Anomalies**  $\neq$  well-modeled points with sensitive behavior!

- 1 Training with a **robust loss** (e.g., Cauchy) to detect badly reconstructed points.
- 2 Use physics knowledge to determine anomalies among them.
- 3 New training from scratch with a **masked non-robust loss function** (e.g., MSE), **ignoring the abnormal outputs**.

## Proposition 3: outputs clustering

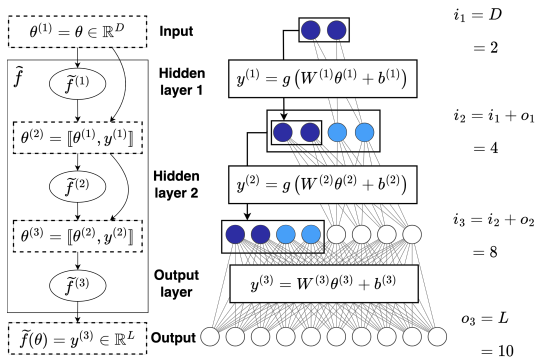
We split lines into clusters based on their **similarity**, and train **dedicated networks** for each cluster.

Method: Spectral clustering.

# Evolutions of a standard multilayer perceptron

## Proposition 4: reuse intermediate computations

As some outputs can be computed from other outputs, keeping track of **intermediate results** optimizes network capacities.



Dense architecture with a growing factor of 2

# Results on the Meudon PDR code

We compare the results obtained with the neural network with interpolation methods.

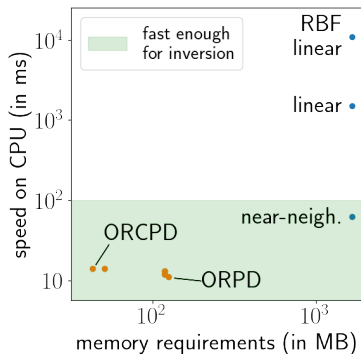
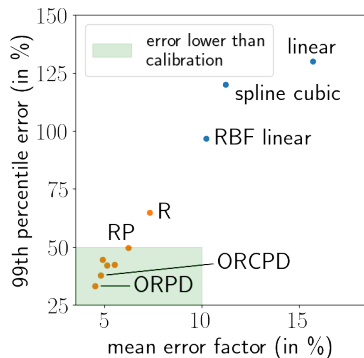
## Metrics

- **Accuracy** – error factor, a kind of symmetrized relative error, computed on **unseen data** (testing set).
- **Memory** – size of the model (whole grid for interpolations)
- **Speed** – computed on a laptop for a batch of 1 000 entries.

To be usable in inference, emulation must satisfy some constraints on these metrics.



# Results on the Meudon PDR code



Summary of results and comparison with interpolation methods.

**R:** regression by an ANN   **O:** outliers removal   **P:** polynomial expansion

**C:** lines clustering   **D:** dense architecture

# Conclusion

## Take-home messages

- Deep learning is efficient to emulate complex simulations, especially with additional constraints.
- Emulators can be plugged in Bayesian inference.
- AI benefits from physical knowledge and rigorous data analysis.

- **Paper:** Palud, Einig et al., A&A, 677, A158 (2023)
- **GitHub:** `einigl/ism-model-nn-approximation`
- **PyPI:** `pip install nnbma`

# References



J. Pety et al. (2017)

The anatomy of the Orion B giant molecular cloud: A local template for studies of nearby galaxies

*Astronomy & Astrophysics*, 599, p. A98



A. Licciardi A and J. Chanussot (2015)

Nonlinear PCA for visible and thermal hyperspectral images quality enhancement

*IEEE Geoscience and Remote Sensing Letters*, 12, pp. 1228-1231



A. Marchal et al. (2019)

ROHSA: Regularized Optimization for Hyper-Spectral Analysis-Application to phase separation of 21 cm data

*Astronomy & Astrophysics*, 626, A101